## EASTERN UNIVERSITY, SRI LANKA

SPECIAL DEGREE EXAMINATION IN MATHEMATICS
ACADEMIC YEAR - 2013/2014 (June, 2016)

## MTS 08 - RELATIVITY

1. In two spacetime dimensions two observers moving with constant relative velocity $v$ set up coordinate system $(c t, x)$ and $\left(c t^{\prime}, x^{\prime}\right)$ respectively. Show that if they set their clocks to $t=t^{\prime}=0$ when pass each other, the transform between these coordinate systems is the Lorenz transform:

$$
\binom{c t}{x}=\gamma\left(\begin{array}{cc}
1 & \frac{v}{c} \\
\frac{v}{c} & 1
\end{array}\right)\binom{c t^{\prime}}{x^{\prime}}, \text { where } \gamma=\left(1-\frac{v^{2}}{c^{2}}\right)^{-\frac{1}{2}} .
$$

Let $K$ and $K^{\prime}$ be two inertial frames such that the origin of $K^{\prime}$ moves with relative speed $v$ in the $x$-direction.
(a) In $K$ a rod at rest has length $l_{0}$. What is the length of the rod in $K^{\prime}$ ?
(b) Let $A$ and $B$ be two simultaneous events in $K$ and suppose $A$ is at $(0,0)$ and $B$ is at $(0, x)$ where $x \neq 0$. Show that $A$ and $B$ are not simultaneous in $K^{\prime}$.
(c) Show that a particle moving with the speed of light along the $x$-direction in $K$ also moves at the speed of light in $K^{\prime}$.
2. (a) Two particles with rest mass $m_{1}$ and $m_{2}$ are moving along $x$-direction with velocities $u_{1}$ and $u_{2}$. These particles collides and form a new particle with rest mass $m_{3}$ and also moving along the $x$-direction with velocity $u_{3}$. Show that

$$
\begin{aligned}
& \qquad u_{3}=\frac{m_{1} \gamma_{1} u_{1}+m_{2} \gamma_{2} u_{2}}{m_{1} \gamma_{1}+m_{2} \gamma_{2}} \text { and } m_{3}^{2}=m_{1}^{2}+m_{2}^{2}+2 m_{1} m_{2} \gamma_{1} \gamma_{2}\left(1-\frac{u_{1} u_{2}}{c^{2}}\right) \\
& \text { where } \gamma_{i}=\gamma\left(u_{i}\right)=\left(1-\frac{u_{i}^{2}}{c^{2}}\right)^{-1 / 2} .
\end{aligned}
$$

(b) Suppose that a photon is traveling along $x$-axis and collides with a statio electron of mass $m$. After the collision the photon and electron move in plane and making angles of $\theta$ (anti-clockwise) and $\phi$ (clockwise) with $x$ respectively. Show that:
i.

$$
\bar{\nu}=\frac{\nu}{1+\left(\frac{h \nu}{m c^{2}}\right)(1-\cos \theta)},
$$

where $\nu$ and $\bar{\nu}$ are the frequencies of the photon before and after collision and $h$ is the Plank's constant.
ii.

$$
\sin ^{2} \frac{\theta}{2}=\frac{m c}{2}\left(\frac{1}{q}-\frac{1}{p}\right)
$$

where $p$ and $q$ are the momentum of photon before and after the col
3. (a) Show that the Riemann tensor

$$
R^{d}{ }_{a b c}=\Gamma^{d}{ }_{a c, b}-\Gamma^{d}{ }_{a b, c}+\Gamma^{e}{ }_{a c} \Gamma^{d}{ }_{e b}-\Gamma^{e}{ }_{a b} \Gamma^{d}{ }_{e c}
$$

arises from the equation $V_{a ; b c}-V_{a ; c b}=R^{d}{ }_{a b c} V_{d}$.
(b) Using the Bianchi identity

$$
R_{b c d ; e}^{a}+R_{b d e ; c}^{a}+R_{b e c ; d}^{a}=0
$$

show that $G^{a b} ; b=0$.
(c) Prove the following:
i. $R_{b c d}^{a}+R_{c d b}^{a}+R_{d b c}^{a}=0$;
ii. $\lambda_{a ; b c}=R_{a b c d} \lambda^{d}$ if $\lambda_{a ; b}+\lambda_{b ; a}=0$.
4. Use the Euler-Lagrange equations to obtain the non-vanishing Christoffel for the metric

$$
d s^{2}=-A(r) d t^{2}+B(r) d r^{2}+r^{2}\left(d \theta^{2}+\sin ^{2} \theta d \phi^{2}\right)
$$

Hence show that the only non-zero Ricci tensor components for this metric a by

$$
\begin{aligned}
& R_{00}=-\frac{A^{\prime \prime}}{2 B}+\frac{A^{\prime}}{4 B}\left(\frac{A^{\prime}}{A}+\frac{B^{\prime}}{B}\right)-\frac{A^{\prime}}{r B} \\
& R_{11}=\frac{A^{\prime \prime}}{2 A}-\frac{A^{\prime}}{4 A}\left(\frac{A^{\prime}}{A}+\frac{B^{\prime}}{B}\right)-\frac{B^{\prime}}{r B} \\
& R_{22}=\frac{1}{B}-1+\frac{r}{2 B}\left(\frac{A^{\prime}}{A}-\frac{B^{\prime}}{B}\right) \\
& R_{33}=R_{22} \sin ^{2} \theta
\end{aligned}
$$

where $\left(x^{0}, x^{1}, x^{2}, x^{3}\right)=(t, r, \theta, \phi)$ and prime denotes the differentiation with respect to $r$.
5. (a) Using the result obtained in question 4, derive the vacuum solution of the Einstein field equations for the static, exterior geometry of a massive object

$$
d s^{2}=-c^{2}\left(1-\frac{2 m}{r}\right) d t^{2}+\left(1-\frac{2 m}{r}\right)^{-1} d r^{2}+r^{2}\left(d \theta^{2}+\sin ^{2} \theta d \phi^{2}\right)
$$

clearly stating all results used
(b) Let the equations for a particle be

$$
\begin{aligned}
\left(1-\frac{2 m}{r}\right) \dot{t} & =k \\
r^{2} \dot{\phi} & =h \\
c^{2}\left(1-\frac{2 m}{r}\right) \dot{t}^{2}-\left(1-\frac{2 m}{r}\right)^{-1} \dot{r}^{2}-r^{2} \dot{\phi}^{2} & =c^{2}
\end{aligned}
$$

Show that the following results hold in vertical free fall:

$$
\begin{aligned}
k & =\sqrt{1-2 m / r_{0}} \\
\ddot{r}+\frac{m c^{2}}{r^{2}} & =0 \\
\frac{1}{2} \dot{r}^{2} & =m c^{2}\left(\frac{1}{r}-\frac{1}{r_{0}}\right)
\end{aligned}
$$

where $r_{0}$ is the point of release of the particle.
6. (a) Let $u$ be an affine parameter along the null geodesic of the light pulse. Show that the change in coordinate time can be written as

$$
\frac{d t}{d u}=\frac{1}{c}\left[\left(1-\frac{2 m}{r}\right)^{-1} g_{i j} \frac{d x^{i}}{d u} \frac{d x^{j}}{d u}\right]^{1 / 2}
$$

which integrates to give

$$
t_{R}-t_{E}=\frac{1}{c} \int_{u_{E}}^{u_{R}}\left[\left(1-\frac{2 m}{r}\right)^{-1} g_{i j} \frac{d x^{i}}{d u} \frac{d x^{j}}{d u}\right]^{1 / 2} d u
$$

where $E$ and $R$ denotes the emitter and receiver respectively.
(b) Using the expression above explain why the change in coordinate time between the time of emission between consecutive signals $\left(\Delta t_{E}\right)$ is the same as the change in coordinate time between time of reception $\left(\Delta t_{R}\right)$ of these signals? Using this fact, argue that the change in proper time between emission and reception is given by

$$
\frac{\Delta \tau_{R}}{\Delta \tau_{E}}=\left(\frac{1-\frac{2 m}{\tau_{R}}}{1-\frac{2 m}{r_{E}}}\right)^{1 / 2}
$$

(c) If the pulses are emitted and received with frequencies $\nu_{e}=n / \Delta \tau_{E}$ and $n / \Delta \tau_{R}$, show that in the limit of $r / m$ being small the spectral shift is givel

$$
\frac{\Delta \nu}{\nu_{E}} \equiv \frac{\nu_{R}-\nu_{E}}{\nu_{E}} \approx \frac{G M}{c^{2}}\left(\frac{1}{r_{R}}-\frac{1}{r_{E}}\right)
$$

(d) The wavelength of a helium-neon lases is measured inside a Skylab freely ing far out in deep space, and is found to be 632.8 nm . What wavelength an experimenter measure if:
i. he and the laser fell freely together towards a neutron star?
ii. he remained in the freely floating Skylab while the laser transmitted ra from the surface of the neutron star of mass $10^{30} \mathrm{~kg}$ and radius $r_{B}=$ iii. he were beside the laser, both on the surface of the neutron star? iv. he were on the surface of neutron star while the laser was back distant Skylab?

