EASTERN UNIVERSITY, SRI LANKA THIRD EXAMINATION IN SCIENCE - 2008/2009 SECOND SEMESTER (PROPER/REPEAT) (October/November 2010) PH 305 FUNDAMENTALS OF STATISTICAL PHYSICS

Time: 01 hour.

Answer <u>ALL</u> Questions

1. Derive Maxwell-Boltzmann distribution function in statistical physics, using the thermodynamic probability distribution function; IBRARY

$$\Omega = N! \prod_{j} \frac{g_{j}^{n_{j}}}{n_{j}!} \text{ and } \beta = \frac{1}{KT} \left(1 \text{ IFER } 2 \right)$$

Consider a gas of N magnetic atoms with spin $\frac{1}{2}$ due to one unpaired electrons. Each atom has intrinsic magnetic moment μ and the interaction between atoms is negligible. The magnetic moment of an atom is: $\mu = \mp \frac{1}{2}g_{s}\mu_{B}$

- Where: μ Magnetic moment
 - g_s Degeneracy of a state
 - μ_B Magnetic moment due to magnetic field
- a) Write down expressions for energies correspond to Spin Down (μ anti-parallel to B) and Spin Up (μ parallel to B).
- b) Express Maxwell-Boltzmann distribution function for this system.
- c) Obtain an expression for partition function for this system.
- d) Determine the total energy of the system by using above expressions.
- e) Hence, show that the magnetization of the electron is:

$$M = -rac{E}{BV}$$
 where $M = rac{1}{V} \sum n_j \mu_j$

 State the conditions under which a system of particles obeys the Bose-Einstein distribution law and derive an expression for the corresponding distribution.

You may use the thermodynamic probability distribution function;

$$\Omega = \prod_{j} \frac{(m_j + g_j - 1)!}{m_j! (g_j - 1)!} \text{ and } \beta = \frac{1}{KT}$$

Consider a quantum mechanical gas of non interacting spin zero bosons each of mass m which is free to move within the volume V at a temperature T. The Density of State of the boson is:

$$D(\omega) = \frac{V\omega^2}{\pi^2 C^3} d\omega$$

- a) Obtain an expression for energy of bosons U(T) in the region.
- b) Show how the expression obtained in part (a) is modified for photon (mass = 0) gas in the distribution function at very low temperature as $\mu \rightarrow 0$.
- c) Hence prove that the energy U(T) satisfied the black body radiation using the function:

$$\int_0^\infty \frac{x^3}{e^x - 1} \, dx = \frac{\pi^4}{15}$$